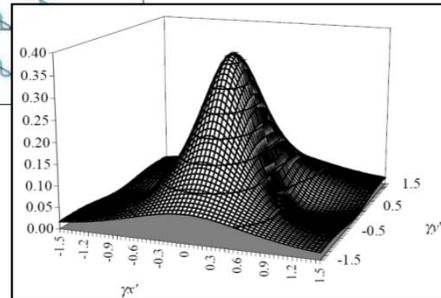
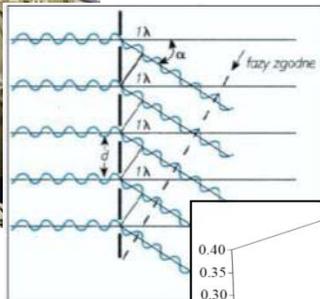
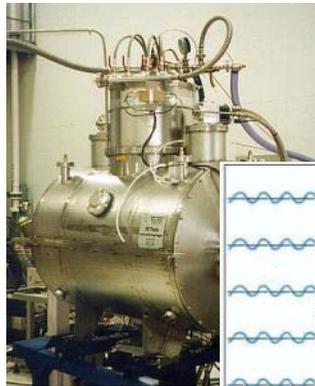


# Insertion Devices

## Lecture 2

### Wigglers and Undulators



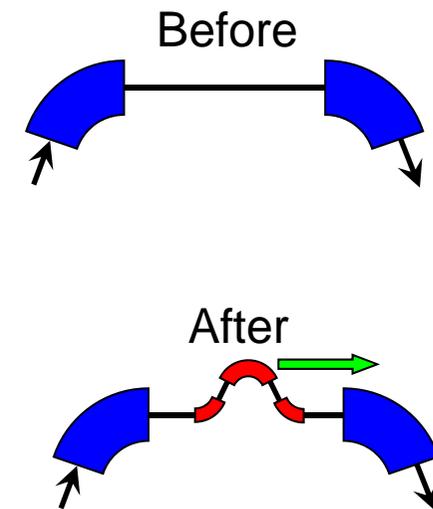
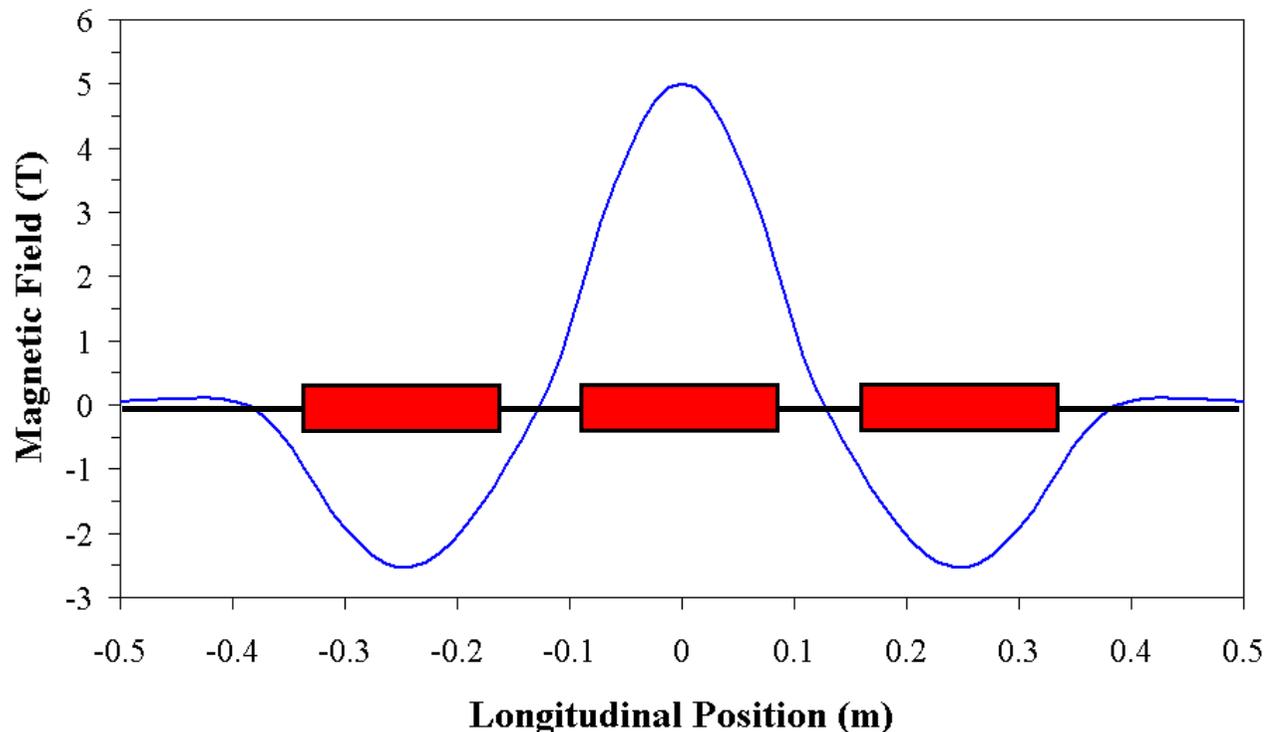
Jim Clarke  
ASTeC  
Daresbury Laboratory

# Wavelength Shifters

How can you put a high magnetic field into a ring?

A popular solution is to use 3 magnets to create a chicane-like trajectory on the electron beam in a straight section

The central magnet is the **high field** bending magnet source

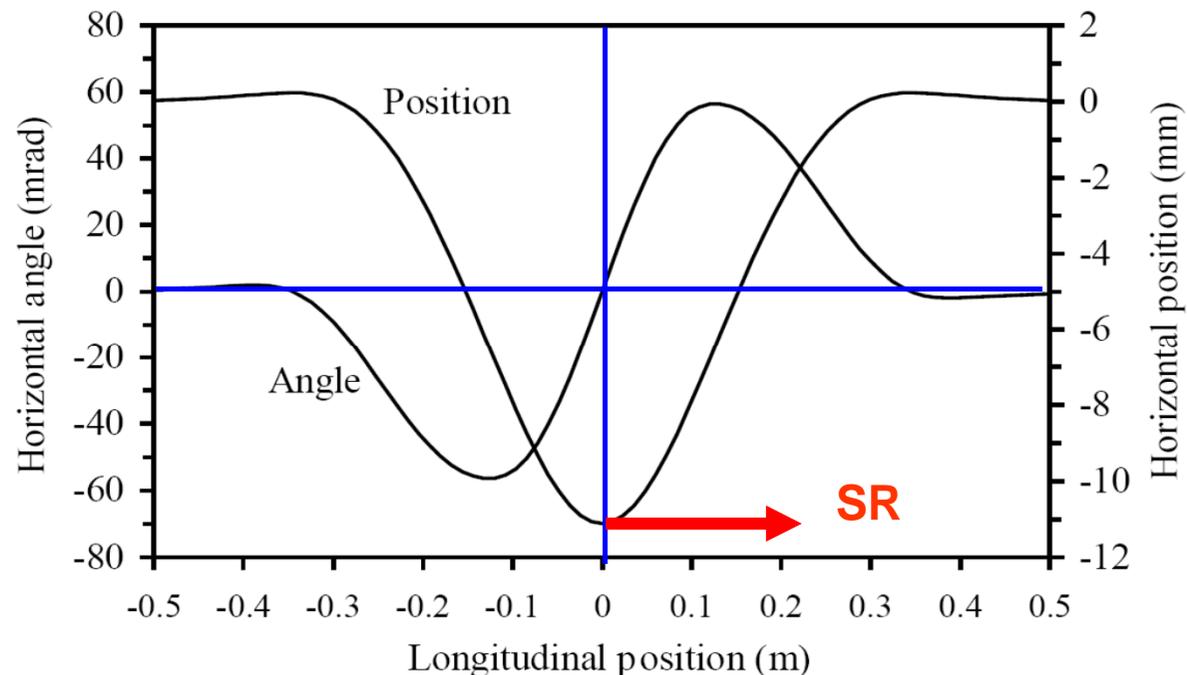


## Electron trajectory in a Wavelength Shifter

The electron enters on axis and exits on axis (“**Insertion Device**”)

The peak of the trajectory bump occurs at the peak magnetic field – when the angle is zero

SR emitted here will travel parallel to the beam axis (at a tangent to the trajectory)



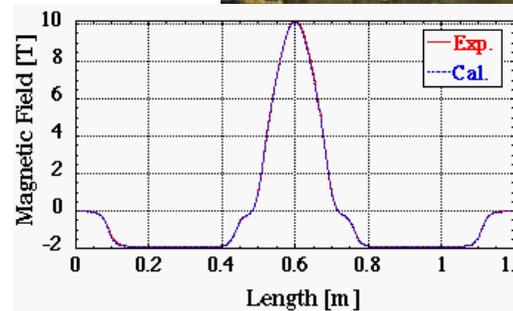
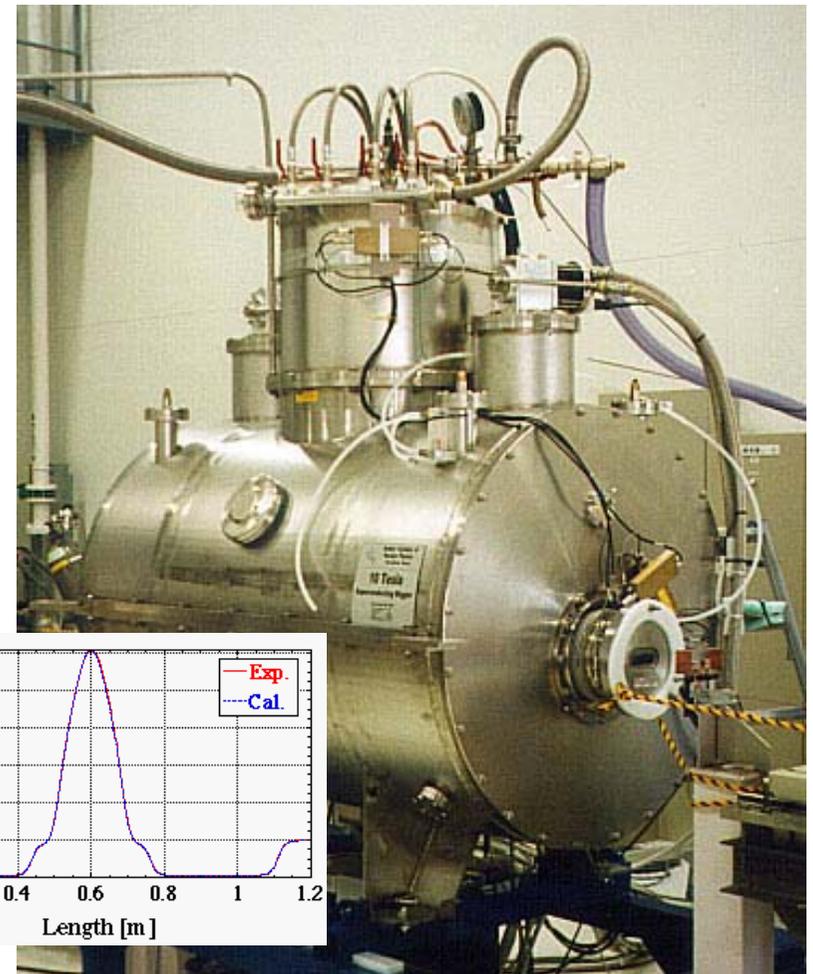
# Examples of Wavelength Shifters



SRS 6T (central pole) wavelength shifter

Wavelength shifters are always superconducting magnets

Spring-8 10T wiggler

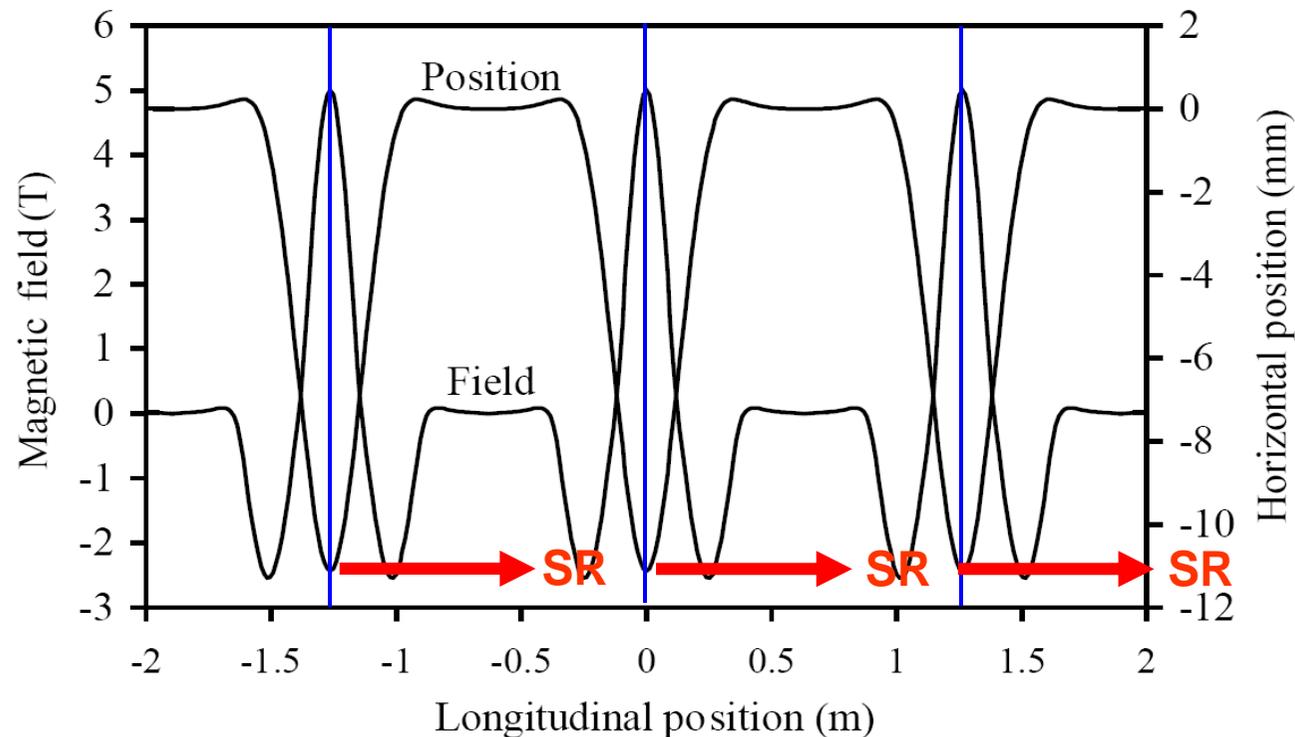


## Extension to Multipole Wigglers

**One** wavelength shifter will give **enhanced flux at high photon energies**

SR is emitted parallel to the axis at the peak of the main pole

Imagine many WS installed next to each other in the same straight ...



## Multiple Wavelength Shifters

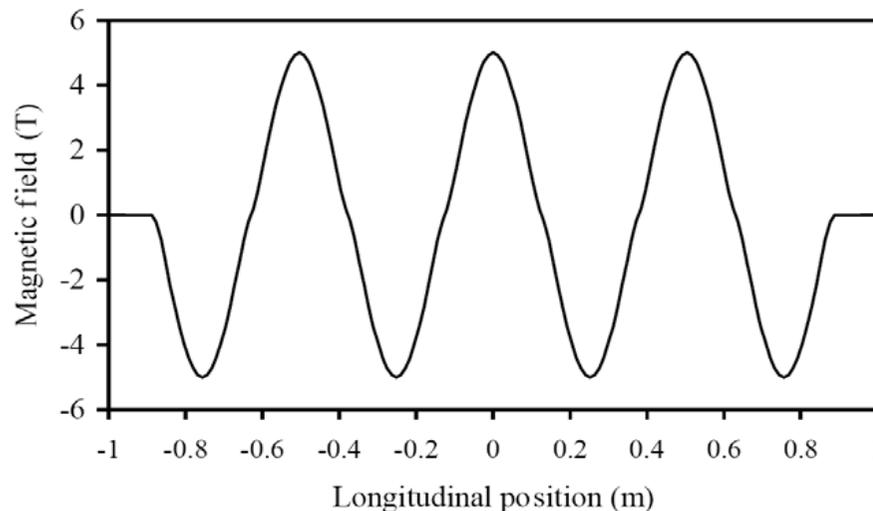
Each WS would be an independent source of SR – all emitting in the forward direction.

The observer (on-axis) would see SR from all 3 Source points

**The observer will therefore see 3 times more flux**

This is the basic concept for a **multipole wiggler**

Three separate WS is not the most efficient use of the space!  
A better way of packing more high field emitters into a straight is...



B field is usually close to sinusoidal

## Multipole Wigglers – Electron Trajectory

---

Electrons travelling in the  $s$  direction

Assuming small angular deflections ( $\dot{x} \ll 1, \dot{y} \ll 1$ )

The equations of motion for the electron are

$$\ddot{x} = \frac{d^2x}{ds^2} = \frac{e}{\gamma m_0 c} (B_y - \dot{y} B_s)$$

$$\ddot{y} = \frac{d^2y}{ds^2} = \frac{e}{\gamma m_0 c} (\dot{x} B_s - B_x)$$

If we have a MPW which only deflects in the horizontal plane  
(x) - **only has vertical fields ( $B_y$ ) on axis**

$$\ddot{x} = \frac{eB_y}{\gamma m_0 c}$$
$$\ddot{y} = 0 .$$

## Angular Deflection

The B field is assumed to be **sinusoidal** with period  $\lambda_u$

$$B_y(s) = -B_0 \sin\left(\frac{2\pi s}{\lambda_u}\right)$$

Integrate once to find  $\dot{x}$  which is the horizontal angular deflection from the s axis

$$\dot{x}(s) = \frac{B_0 e}{\gamma m_0 c} \frac{\lambda_u}{2\pi} \cos\left(\frac{2\pi s}{\lambda_u}\right)$$

Therefore, the peak angular deflection is  $\frac{B_0 e}{\gamma m_0 c} \frac{\lambda_u}{2\pi}$

Define the **deflection parameter**

$$K = \frac{B_0 e}{m_0 c} \frac{\lambda_u}{2\pi} = 93.36 B_0 \lambda_u$$

( $B_0$  in T,  $\lambda_u$  in m)

## Trajectory

---

One more integration gives

$$x(s) = \frac{K}{\gamma} \frac{\lambda_u}{2\pi} \sin\left(\frac{2\pi s}{\lambda_u}\right)$$

The peak angular deflection is  $\frac{K}{\gamma}$

Remember that SR is emitted with a typical angle of  $\sim 1/\gamma$

So if  $K < 1$  the electron trajectory will overlap with the emitted cone of SR (**an undulator**)

If  $K \gg 1$  there will be little overlap and the source points are effectively independent – this is the case for a **MPW**

**The boundary between an undulator and a MPW is not actually so black and white as this!**

## MPW Flux

MPW can be considered a series of dipoles, one after the other

There are two source points per period

The flux is simply the product of the number of source points and the dipole flux for that critical energy

The MPW has two clear advantages

The critical energy can be set to suit the science need

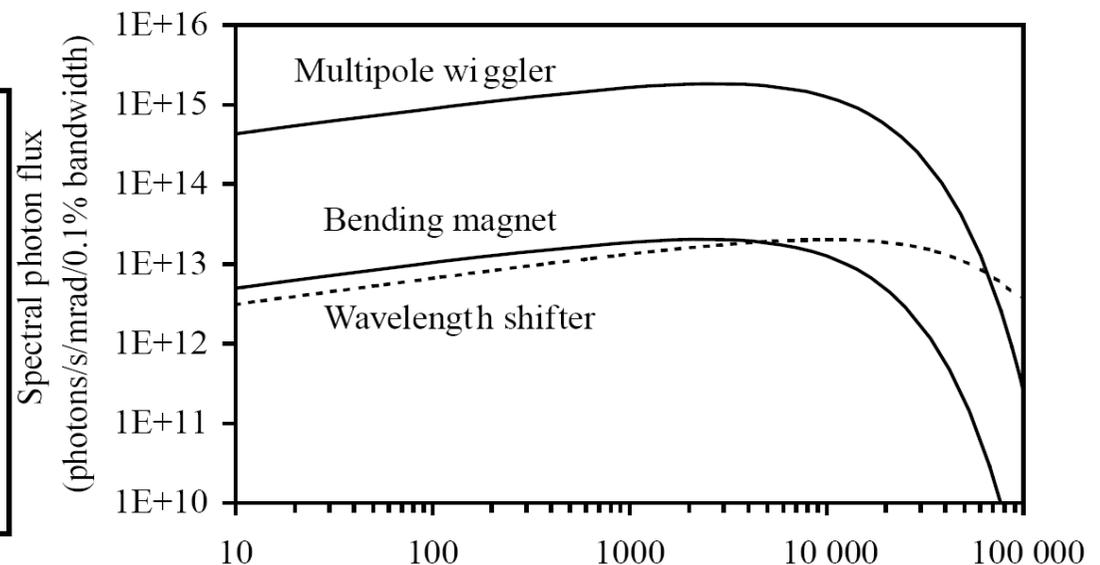
The Flux is enhanced by twice number of periods

300mA, 3 GeV beam

1.4T dipole

6T WS

1.6T, MPW with 45 periods  
(2 poles per period so x90 flux)



## MPW Power

---

The total power emitted by a beam of electrons passing through **any magnet system** is

$$P_{\text{total}} = 1265.5 E^2 I_b \int_0^L B(s)^2 ds$$

**This is a general result – can get the earlier bending magnet result from here.**

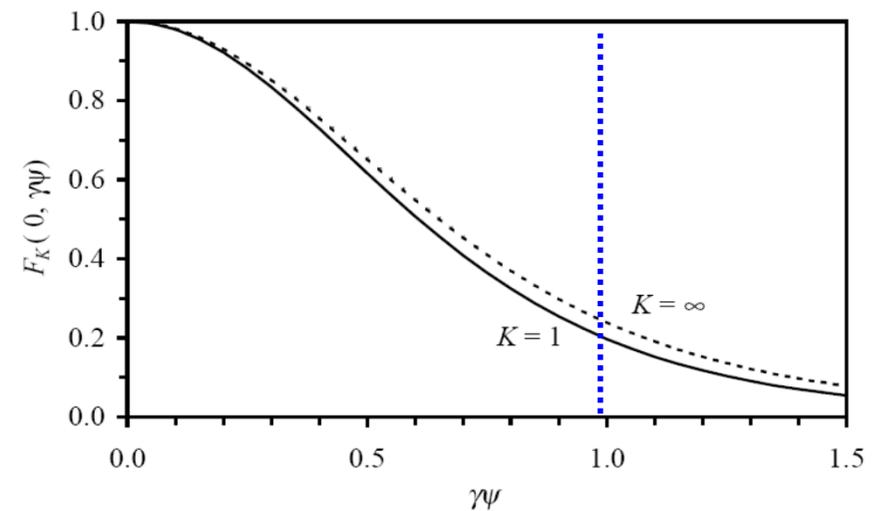
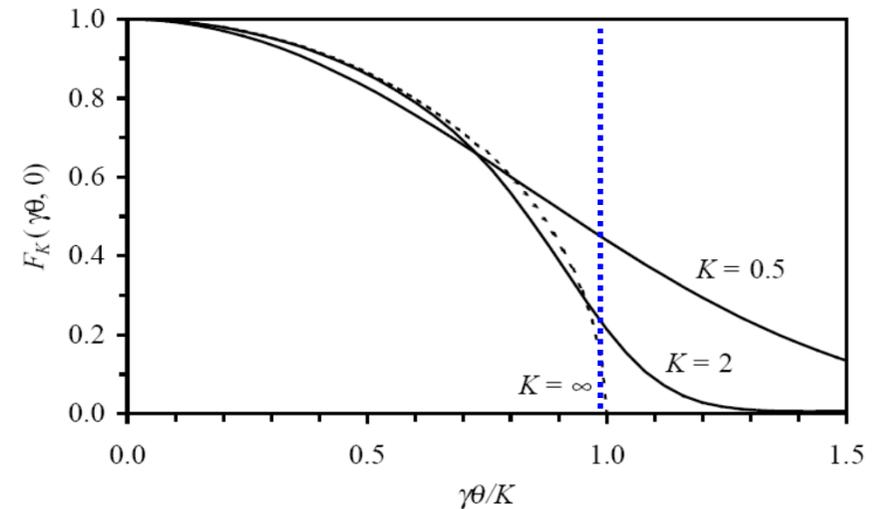
For a sinusoidal magnetic field with peak value  $B_0$  the integral is  $B_0^2 L/2$  and so the total power emitted is (in W)

$$P_{\text{total}} = 632.8 E^2 B_0^2 L I_b$$

# Power Density

The power is contained in  $\pm K/\gamma$  horizontally for large  $K$

Vertically, the power is contained in  $\sim \pm 1/\gamma$



## On-Axis power density

The **Peak** power density is **on-axis**

	Energy (GeV)	$B_0$ (T)	$K$	$L$ (m)	$I_b$ (mA)	$P_{\text{total}}$ (kW)	$dP/d\Omega$ (kW/mrad <sup>2</sup> )
<b>Undulators</b>	2	0.64	3	5	200	1.0	2.2
	3	0.64	3	5	300	3.5	16.8
	6	0.64	3	5	200	9.3	179.1
<b>MPWs</b>	2	2	40	2	200	4.0	0.6
	3	2	40	2	300	13.7	4.9
	6	2	40	2	200	36.4	52.5

## Undulators

For a sinusoidal magnetic field

$$\dot{x}(s) = \frac{dx}{ds} = \frac{K}{\gamma} \cos\left(\frac{2\pi s}{\lambda_u}\right)$$

$\beta_x$  is the relative transverse velocity

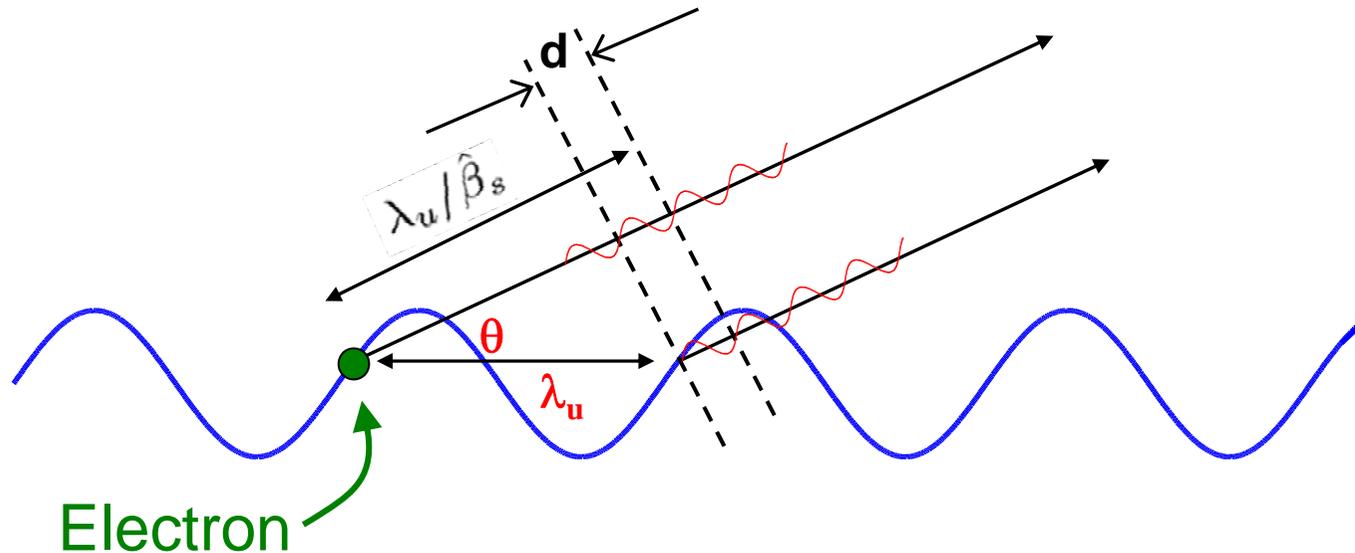
$$\beta_x = \frac{dx/dt}{c} = \frac{K}{\gamma} \cos\left(\frac{2\pi s}{\lambda_u}\right) \quad \Delta s = c\Delta t$$

**The energy is fixed so  $\beta$  is also fixed.** Any variation in  $\beta_x$  will have a corresponding change in  $\beta_s$  ( $\beta_y = 0$ )

$$\begin{aligned}\beta_s^2 &= \beta^2 - \beta_x^2 \\ &= \beta^2 - \frac{K^2}{\gamma^2} \cos^2\left(\frac{2\pi s}{\lambda_u}\right)\end{aligned}$$

# The Condition for Interference

For constructive interference between wavefronts emitted by the same electron **the electron must slip back by a whole number of wavelengths** over one period



Speed = distance/time = c

The time for the electron to travel one period is  $\lambda_u / c\hat{\beta}_s$

In this time the first wavefront will travel the distance  $\lambda_u / \hat{\beta}_s$

## Interference Condition

---

The separation between the wavefronts is  $d = \frac{\lambda_u}{\hat{\beta}_s} - \lambda_u \cos \theta$

And this **must equal a whole number of wavelengths** for constructive interference

$$n\lambda = \frac{\lambda_u}{\hat{\beta}_s} - \lambda_u \cos \theta$$

**n is an integer – the harmonic number**

Using  $(1 - x)^{-1} \sim 1 + x$

and  $1 - \cos \theta = 2 \sin^2(\theta/2)$

and  $\sin \theta \sim \theta$  for small angles:

## Interference Condition

We get

$$n\lambda \sim \frac{\lambda_u \theta^2}{2} + \frac{\lambda_u}{2\gamma^2} + \frac{\lambda_u K^2}{4\beta\gamma^2}$$
$$\sim \frac{\lambda_u}{2\gamma^2} \left( 1 + \frac{K^2}{2} + \theta^2 \gamma^2 \right)$$

And the undulator equation

$$\lambda = \frac{\lambda_u}{2n\gamma^2} \left( 1 + \frac{K^2}{2} + \theta^2 \gamma^2 \right)$$

**Example**, 3GeV electron passing through a 50mm period undulator with  $K = 3$ . First harmonic ( $n = 1$ ), on-axis is  $\sim 4$  nm. **cm** periods translate to **nm** wavelengths because of the huge  $\gamma^2$  term

## Undulator equation implications

---

$$\lambda = \frac{\lambda_u}{2n\gamma^2} \left( 1 + \frac{K^2}{2} + \theta^2 \gamma^2 \right)$$

The wavelength primarily depends on the **period** and the **energy** but also on  $K$  and the observation angle  $\theta$ .

$$K = \frac{B_0 e}{m_0 c} \frac{\lambda_u}{2\pi} = 93.36 B_0 \lambda_u$$

**If we change B we can change  $\lambda$ .** For this reason, undulators are built with smoothly adjustable B field. **The amount of the adjustability sets the tuning range of the undulator.**

**Note.** What happens to  $\lambda$  as B increases?

## Undulator equation implications

---

As **B increases** (and so K), the output **wavelength increases** (photon energy decreases).

This *appears* different to bending magnets and wigglers where we increase B so as to produce shorter wavelengths (higher photon energies).

The wavelength changes with  $\theta^2$ , so it **always gets longer as you move away from the axis**

An important consequence of this is that the **beamline aperture choice is important** because it alters the radiation characteristics reaching the observer.

## Harmonic bandwidth

Assume the undulator contains **N periods**

For constructive interference

$$Nn\lambda = \frac{N\lambda_u}{\hat{\beta}_s} - N\lambda_u \cos \theta$$

For destructive interference to first occur

$$Nn\lambda^* + \lambda^* = \frac{N\lambda_u}{\hat{\beta}_s} - N\lambda_u \cos \theta$$

(ray from first source point exactly out of phase with centre one, ray from 2<sup>nd</sup> source point out of phase with centre+1, etc)

$$Nn\lambda = Nn\lambda^* + \lambda^*$$

Range over which there is some emission  $\Delta\lambda = \lambda - \lambda^*$

**Bandwidth** (width of harmonic line):

$$\frac{\Delta\lambda}{\lambda} \sim \frac{1}{Nn}$$

## Angular width

---

Destructive interference will first occur when

$$Nn\lambda + \lambda = \frac{N\lambda_u}{\hat{\beta}_s} - N\lambda_u \cos \theta^*$$

This gives  $N\lambda_u \cos \theta^* + \lambda = N\lambda_u \cos \theta$

And using  $\cos \theta \sim 1 - \theta^2/2$

We find, for the radiation emitted on-axis, the intensity falls to zero at

$$\Delta\theta = \sqrt{\frac{2\lambda}{N\lambda_u}}$$

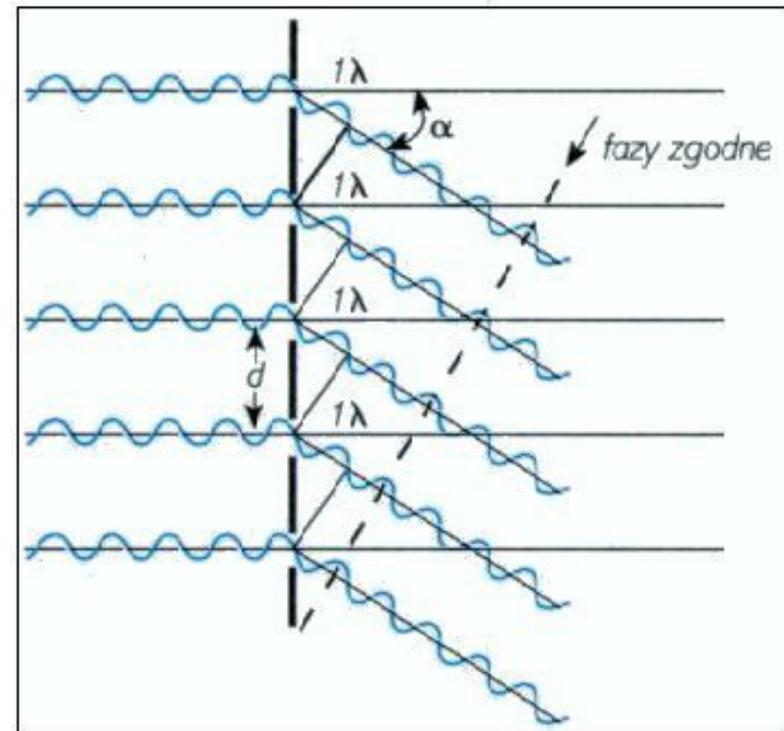
Example, 50mm period undulator with 100 periods emitting 4nm will have  $\Delta\theta = \mathbf{40 \mu rad}$ , significantly less than

$$1/\gamma \sim 170 \mu rad$$

## Diffraction Gratings

Very similar results for **angular width** and **bandwidth** apply to **diffraction gratings**

This is because the diffraction grating acts as a large number of equally spaced sources – **very similar concept as an undulator** (but no relativistic effects!)



The diffraction grating and spectrum on screen  
 $d$  grating constant,  $\lambda$  wave length,  $\alpha$  angle of deflection,

## When does an undulator become a wiggler?

As  $K$  increases the number of harmonics increases:

<b>K</b>	<b>Number of Harmonics</b>
1	few
5	10s
10	100s
20	1000s

At high frequencies the **spectrum smoothes out** and takes on the form of the **bending magnet spectrum**

At low frequencies distinct harmonics are still present and visible

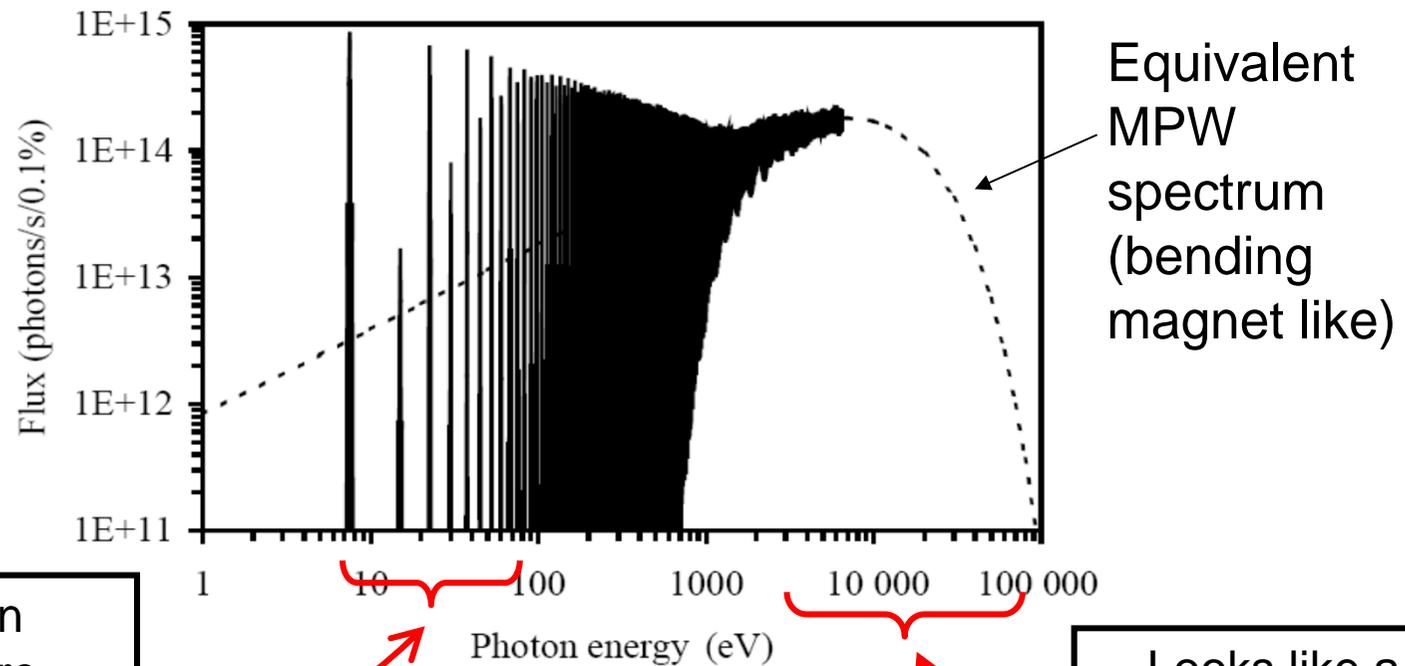
There is in fact **no clear distinction** between an undulator and a wiggler – ***it depends which bit of the spectrum you are observing***

# Undulator or Wiggler?

The difference depends upon which bit of the spectrum you use!

This example shows an undulator calculation for  $K = 15$ .

[Calculation truncated at high energies as too slow!]



Looks like an undulator here

Looks like a wiggler here

## Angular Flux Density

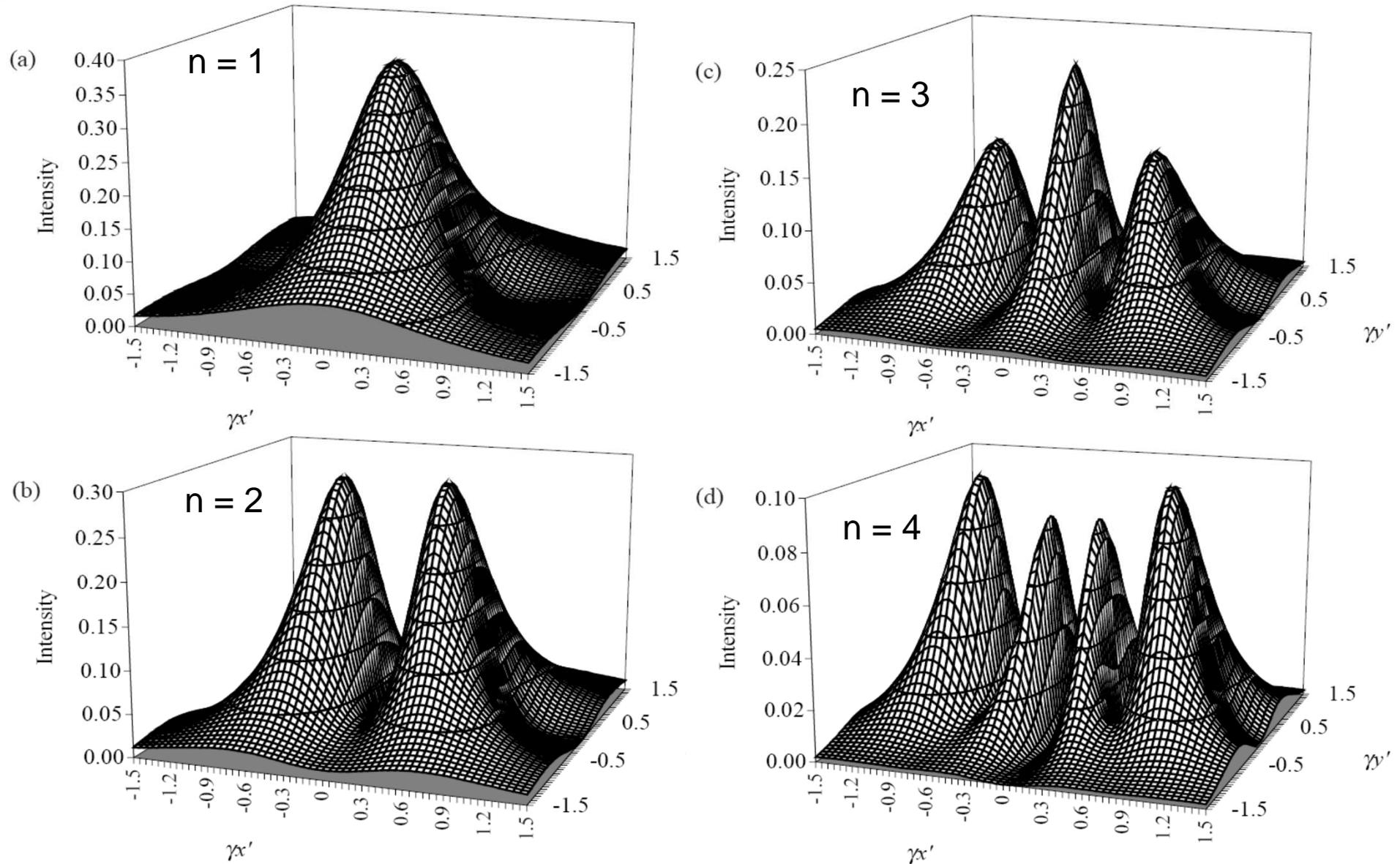
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We want to gain an appreciation for the emitted radiation from an undulator

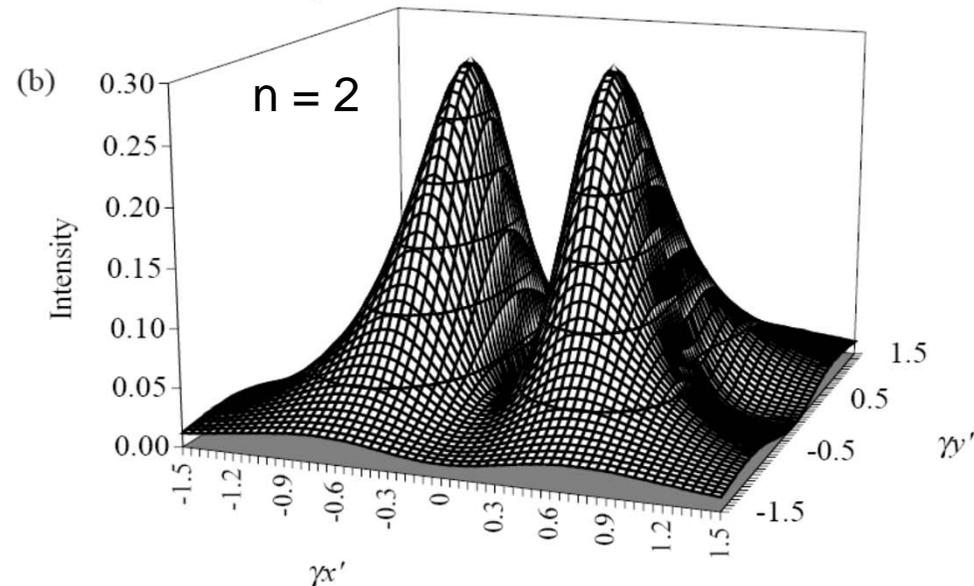
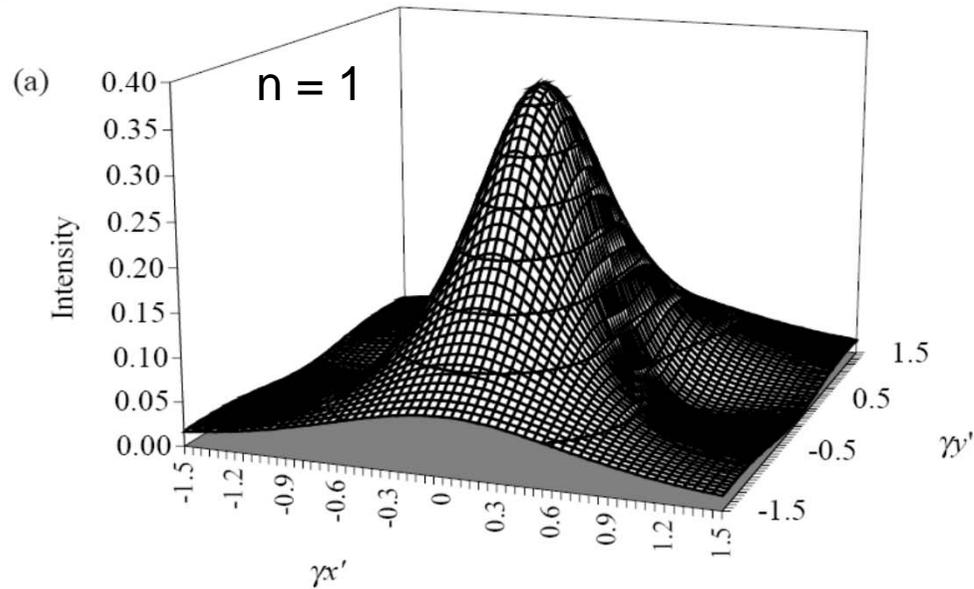
One useful parameter is the (angular) flux density (photons per solid angle) as a function of observation angle

Later we will look at the flux levels and also the polarisation of the radiation

# Angular Flux Density: $K = 1$ examples



## Angular Flux Density: $K = 1$ examples



There are  $n$  peaks in the horizontal plane

The even harmonics have zero intensity on axis

Remember that the wavelength changes with angle ( $\theta^2$  term in the undulator equation) so these plots are **not at a fixed wavelength**

## On Axis (Spectral) Angular Flux Density

In units of photons/sec/mrad<sup>2</sup>/0.1% bandwidth

$$\left. \frac{d\dot{N}}{d\Omega} \right|_{\theta=0} = 1.74 \times 10^{14} N^2 E^2 I_b F_n(K)$$

Where:

N is the number of periods

E is the electron energy in GeV

I<sub>b</sub> is the beam current in A

F<sub>n</sub>(K) is defined below (J are Bessel functions)

$$F_n(K) = \frac{n^2 K^2}{(1 + K^2/2)^2} \left( J_{(n+1)/2}(Y) - J_{(n-1)/2}(Y) \right)^2$$

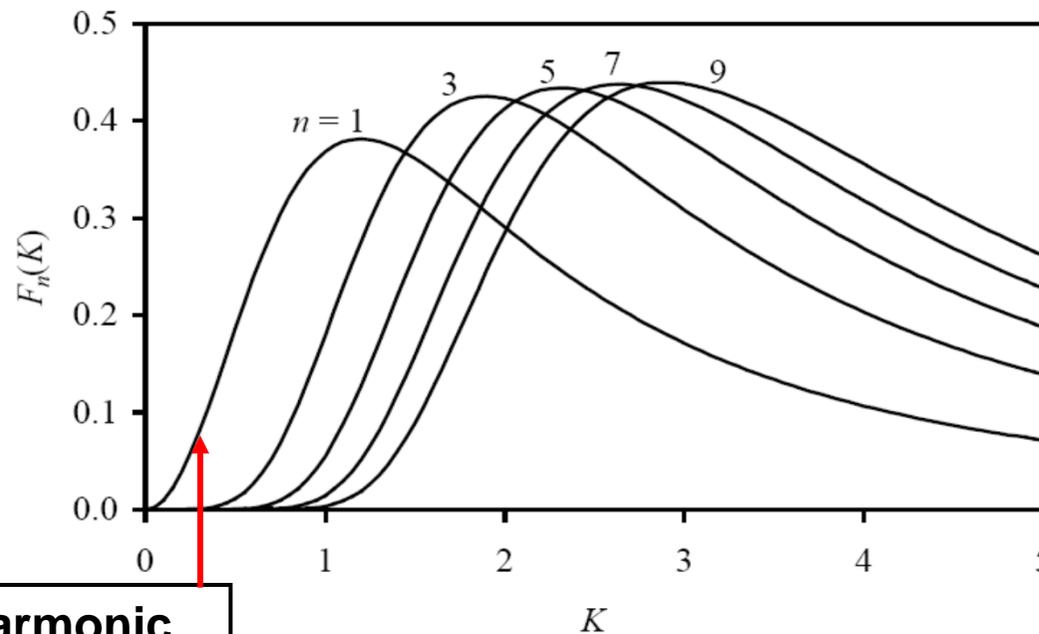
$$Y = \frac{nK^2}{4(1 + K^2/2)}$$

## On Axis (Spectral) Angular Flux Density

In units of photons/sec/mrad<sup>2</sup>/0.1% bandwidth

$$\left. \frac{d\dot{N}}{d\Omega} \right|_{\theta=0} = 1.74 \times 10^{14} N^2 E^2 I_b F_n(K)$$

As  $K$  increases we can see that the influence of the higher harmonics grows



Only 1 harmonic  
at low  $K$

Higher harmonics  
have higher flux  
density

## Example On Axis Angular Flux Density

---

An Undulator with 50mm period and 100 periods with a 3GeV, 300mA electron beam will generate:

Angular flux density of  $8 \times 10^{17}$  photons/sec/mrad<sup>2</sup>/0.1% bw

For a bending magnet with the same electron beam we get a value of  $\sim 5 \times 10^{13}$  photons/sec/mrad<sup>2</sup>/0.1% bw

The undulator has a flux density  $\sim 10,000$  times greater than a bending magnet due to the  $N^2$  term

## Flux on-axis

Previously we argued that light of the same wavelength was contained in a narrow angular width

(interference effect)

$$\Delta\theta = \sqrt{\frac{2\lambda}{N\lambda_u}}$$

Assuming that the SR is emitted in angle with a Gaussian distribution with standard deviation  $\sigma_{r'}$  then we can approximate:

$$\sigma_{r'} = \sqrt{\frac{\lambda}{N\lambda_u}} = \sqrt{\frac{\lambda}{L}}$$

For a Gaussian  $\frac{d\dot{N}}{d\Omega} = \frac{d\dot{N}}{d\Omega} \Big|_{\theta=0} \exp\left(-\frac{\theta^2}{2\sigma_{r'}^2}\right)$

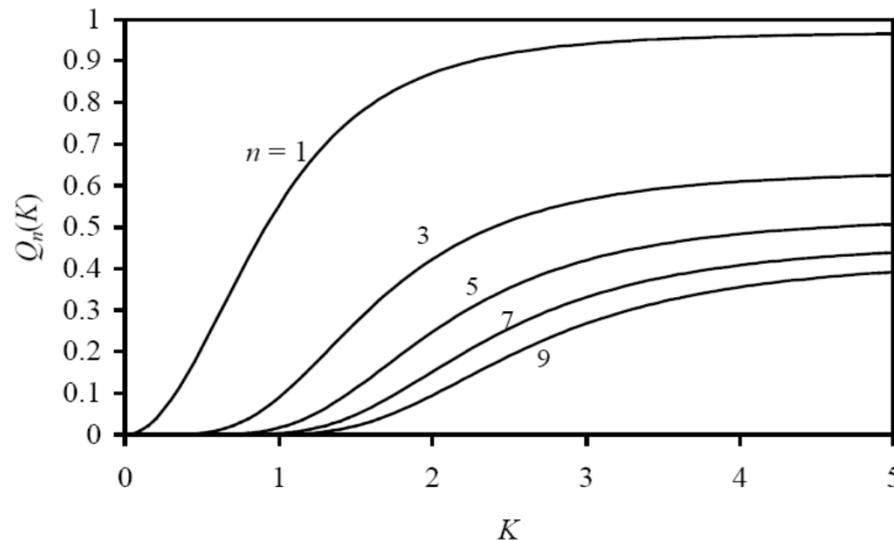
Integrating over all angles gives  $\dot{N} = 2\pi\sigma_{r'}^2 \frac{d\dot{N}}{d\Omega} \Big|_{\theta=0}$

## Flux in the Central Cone

In photons/sec/0.1% bandwidth the flux in this central cone is

$$\dot{N} = 1.43 \times 10^{14} N I_b Q_n(K)$$

$$Q_n(K) = \frac{1 + K^2/2}{n} F_n(K)$$



As  $K$  increases the higher harmonics play a more significant part but the **1st harmonic always has the highest flux**

## Example Flux in the Central Cone

---

Undulator with 50mm period, 100 periods

3GeV, 300mA electron beam

Our example undulator has a flux of  $4 \times 10^{15}$  compared with the bending magnet of  $\sim 10^{13}$

**The factor of few 100 increase is dominated by the number of periods, N (100 here)**

## Undulator Tuning Curves

---

Undulators are often described by “**tuning curves**”

These show the flux (or brightness) *envelope* for an undulator.

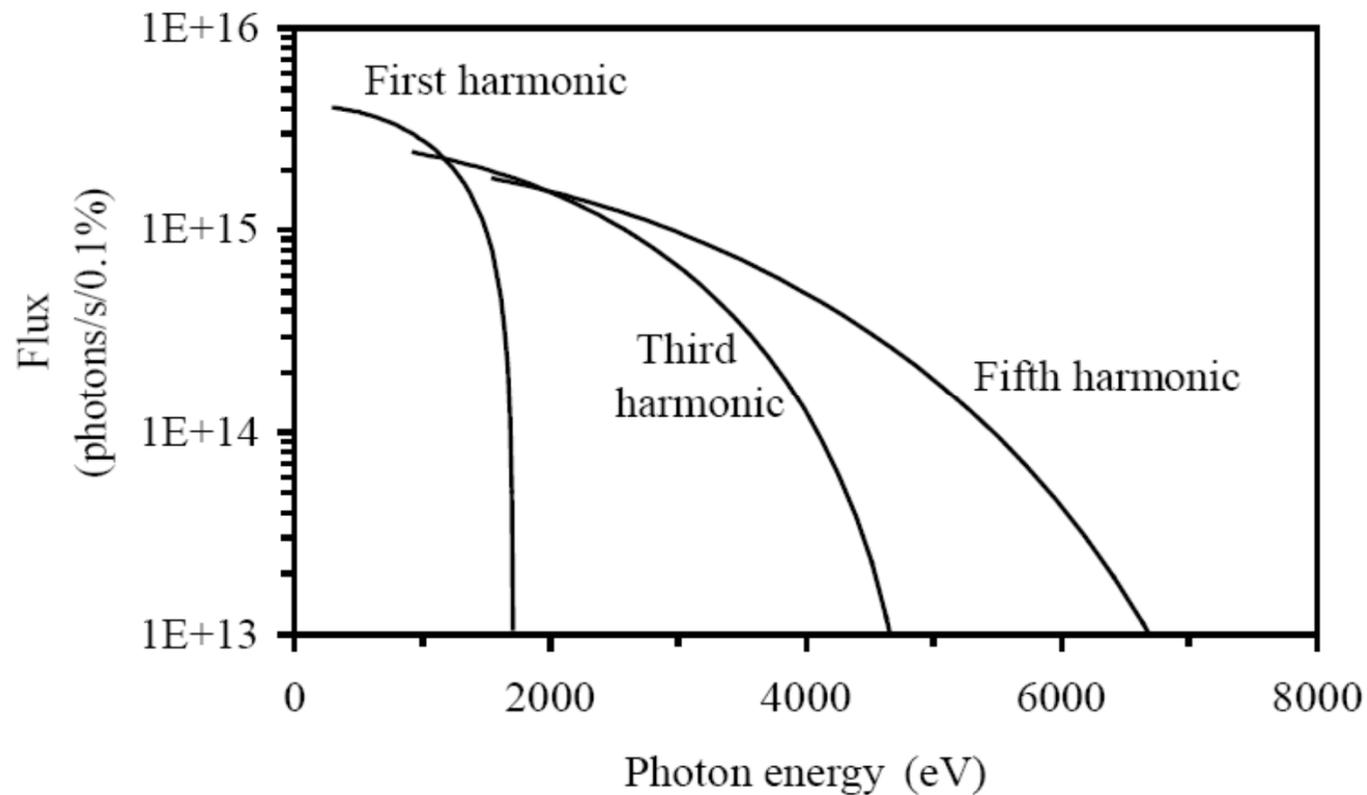
The tuning of the undulator is achieved by **varying the K parameter**

Obviously  $K$  can only be one value at a time !

These curves represent what the undulator is capable of as the  $K$  parameter is varied – but not all of this radiation is available at the same time!

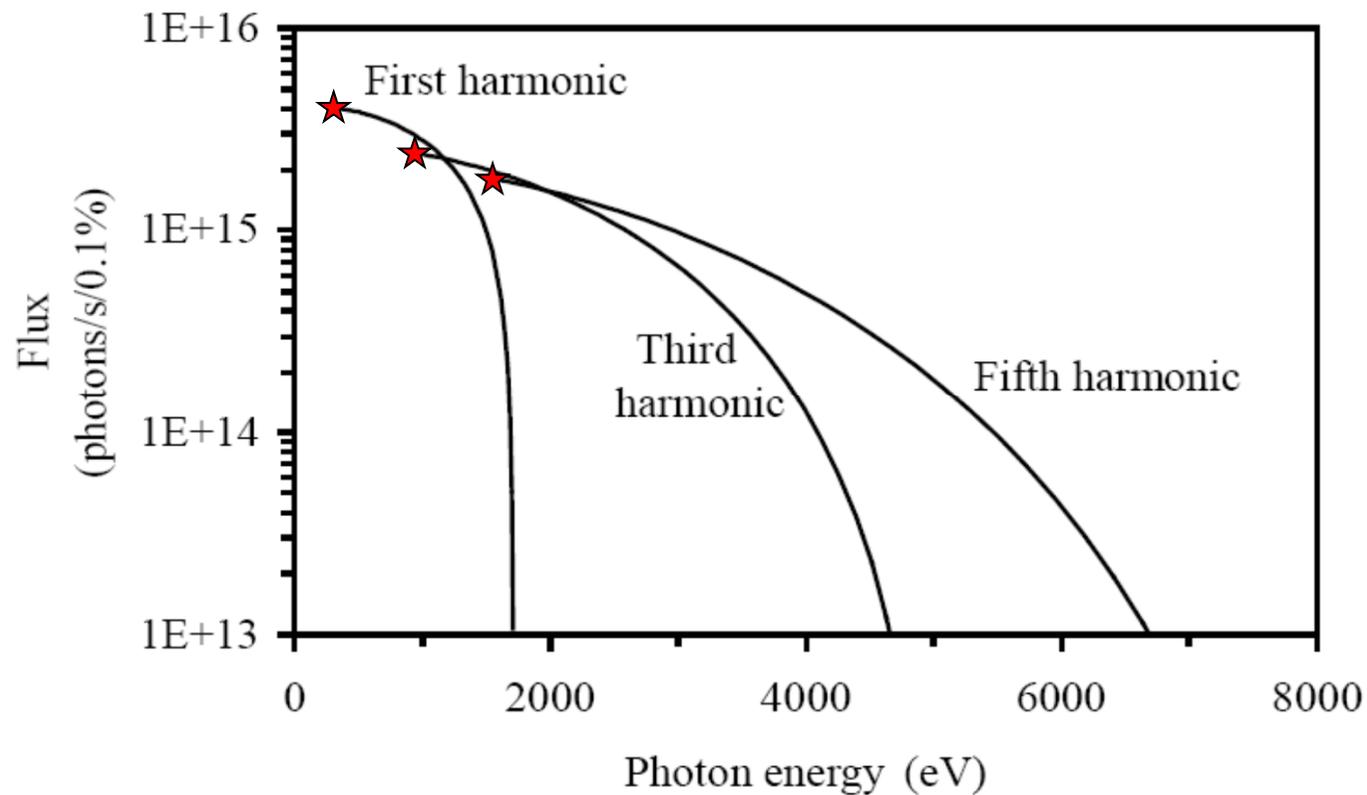
## Example Undulator Tuning Curve

Undulator with 50mm period, 100 periods  
3GeV, 300mA electron beam



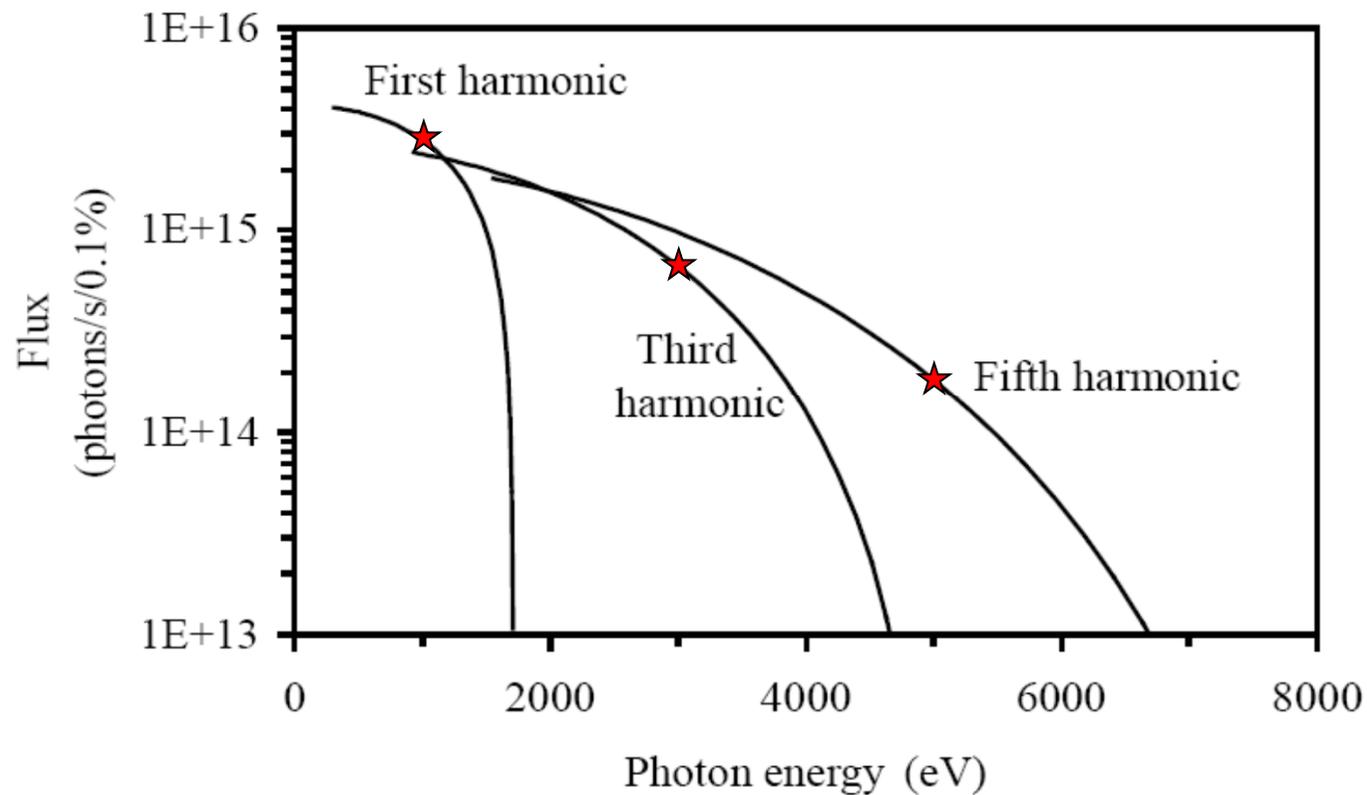
## Example Undulator Tuning Curve

Undulator with 50mm period, 100 periods  
3GeV, 300mA electron beam



## Example Undulator Tuning Curve

Undulator with 50mm period, 100 periods  
3GeV, 300mA electron beam



# Undulator Brightness

All emitted photons have a position and an angle in phase space ( $x, x'$ )

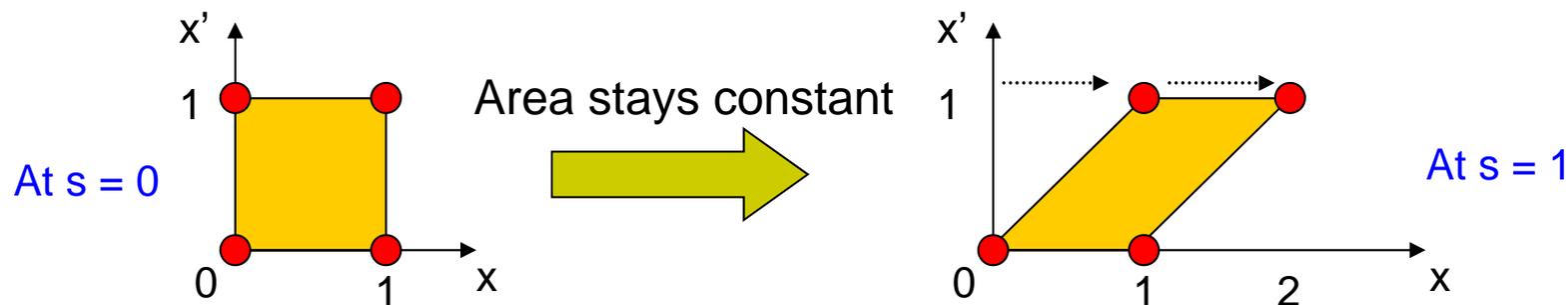
Phase space evolves as photons travel but the area stays constant  
(Liouville's theorem)

The emittance of an electron beam is governed by the same theorem

**Brightness is the phase space density of the flux** – takes account of the number of photons and their concentration

**Brightness (like flux) is conserved** by an ideal optical transport system, unlike angular flux density for instance

Since it is conserved it is a **good figure of merit** for comparing sources (like electron beam emittance)



## Undulator Brightness

---

To calculate the brightness we need the phase space areas

We need to include the photon and electron contributions

We add contributions in quadrature as both are *assumed* to be Gaussian distributions

The photon beam size is found by assuming the source is the fundamental mode of an optical resonator (called the Gaussian laser mode)

$$2\pi\sigma_r\sigma_{r'} = \frac{\lambda}{2}$$

For the peak undulator flux:

$$\sigma_r = \frac{1}{4\pi} \sqrt{\lambda L}$$
$$\sigma_{r'} = \sqrt{\frac{\lambda}{L}} \cdot$$

## Undulator Brightness

---

The example undulator has a source size and divergence of **11 $\mu\text{m}$  and 28 $\mu\text{rad}$**  respectively

The electron beam can also be described by gaussian shape (genuinely so in a storage ring!) and so the **effective** source size and divergence is given by

$$\begin{aligned}\Sigma_x &= \sqrt{\sigma_x^2 + \sigma_r^2} & \Sigma_{x'} &= \sqrt{\sigma_{x'}^2 + \sigma_{r'}^2} \\ \Sigma_y &= \sqrt{\sigma_y^2 + \sigma_r^2} & \Sigma_{y'} &= \sqrt{\sigma_{y'}^2 + \sigma_{r'}^2}\end{aligned}$$

The units of brightness are photons/s/solid area/solid angle/spectral bandwidth

## Example Brightness

---

Undulator brightness is the flux divided by the phase space volume given by these **effective** values

$$B = \frac{\dot{N}}{4\pi^2 \Sigma_x \Sigma_y \Sigma_{x'} \Sigma_{y'}}$$

For our example undulator, using electron beam parameters:

$$\sigma_x = 100 \mu\text{m}, \sigma_y = 10 \mu\text{m}, \sigma_{x'} = 20 \mu\text{rad}, \text{ and } \sigma_{y'} = 2 \mu\text{rad}$$

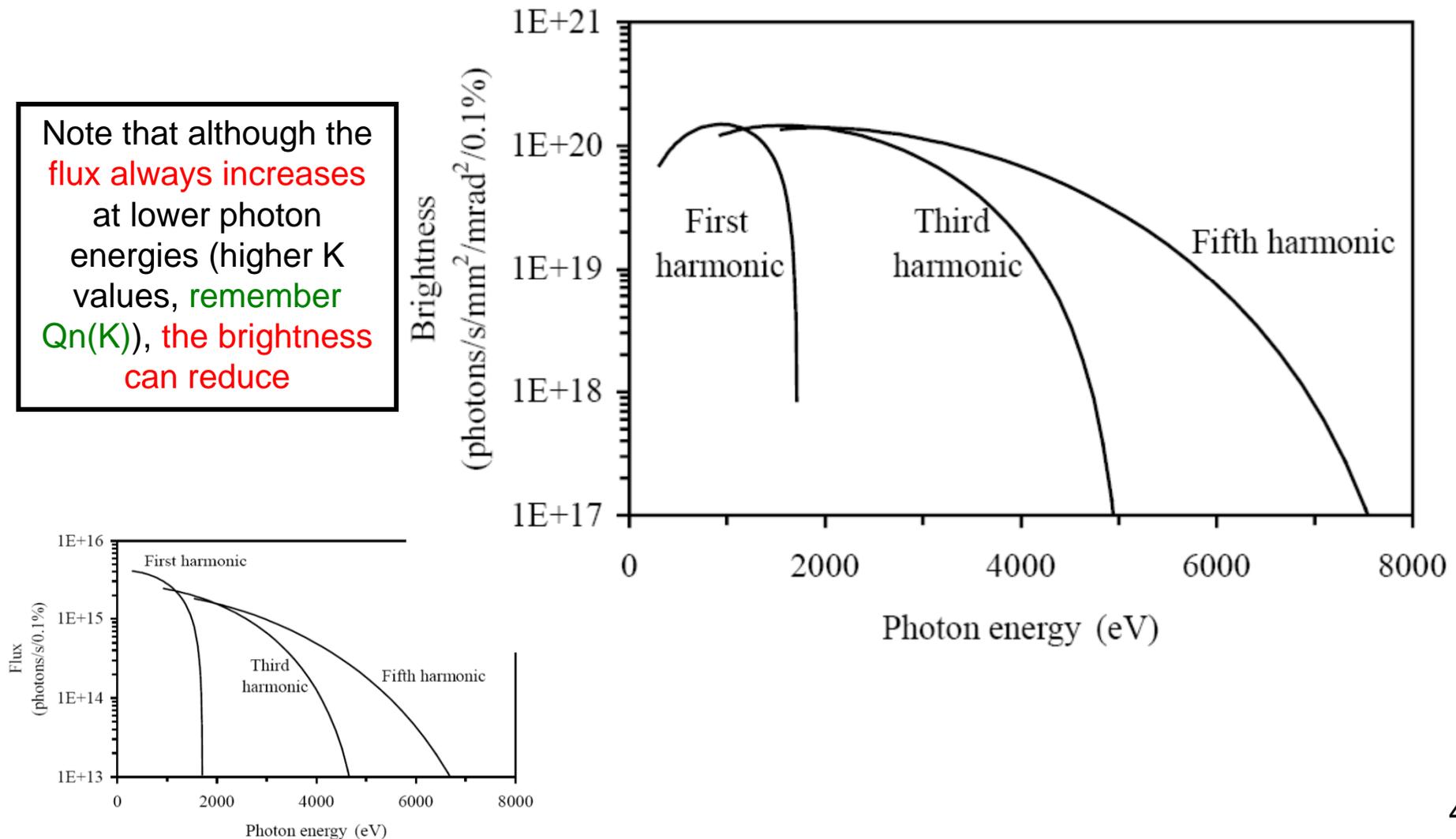
The brightness is

$$7 \times 10^{19} \text{ photons/s/mrad}^2/\text{mm}^2/0.1\% \text{ bandwidth.}$$

# Brightness Tuning Curve

The same concept as the flux tuning curve

Note that although the flux always increases at lower photon energies (higher K values, remember  $Q_n(K)$ ), the brightness can reduce



## Warning!

---

The definition of the photon source size and divergence is somewhat arbitrary

Many alternative expressions are used – for instance we could have used  $\sigma_{r'} = \sqrt{\lambda/2L}$

The actual effect on the absolute brightness levels of these alternatives *is relatively small*

**But** when **comparing your undulator against other sources** it is important to know which expressions have been assumed !

## Summary

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Insertion Devices are added to accelerators to produce light that is specifically tailored to the experimental requirements (wavelength, flux, brightness, polarisation, ...)

Multipole wigglers are periodic, high field devices, used to generate enhanced flux levels (proportional to the number of poles)

Undulators are periodic, low(er) field, devices which generate radiation at specific harmonics

The distinction between undulators and multipole wigglers is not black and white.

Undulator flux scales with  $N$ , flux density with  $N^2$